

Landau Damping

Plasma Physics

Rohit Dokara

JAP, IISc

November 16, 2018

Wave propagation in warm plasmas

- Assumptions
 - Small amplitude
 - Uniform plasma
 - No magnetic field
 - Ions are immobile, only \bar{e} move

Wave propagation in warm plasmas

- Looking for electrostatic plasma waves ($\vec{B} = 0$)
- Vlasov equation for an unmagnetized collisionless plasma

$$\frac{\partial f_e}{\partial t} + \vec{v} \cdot \nabla f_e - \frac{e}{m_e} \vec{E} \cdot \nabla_v f_e = 0$$

where

$f_e(\vec{r}, \vec{v}, t)$ - electron distribution function

$$\vec{E} = -\nabla\phi, \quad \nabla^2\phi = -\frac{\rho}{\epsilon_0} = -\frac{e}{\epsilon_0} \left\{ n - \int f_e d^3\vec{v} \right\}$$

n is the no. density of ions

Vlasov's solution

- Write f_e as equilibrium + perturbation

$$f_e(\vec{r}, \vec{v}, t) = f_0(\vec{v}) + f_1(\vec{r}, \vec{v}, t)$$

- f_0 is the equilibrium distribution and f_1 is the perturbation
- Only for dealing with small amplitude waves
- Note that

$$\int f_0 d^3\vec{v} = n$$

Vlasov's solution

- Using

$$f_e(\vec{r}, \vec{v}, t) = f_0(\vec{v}) + f_1(\vec{r}, \vec{v}, t)$$

and

$$\frac{\partial f_e}{\partial t} + \vec{v} \cdot \nabla f_e - \frac{e}{m_e} \vec{E} \cdot \nabla_{\vec{v}} f_e = 0$$

- We get

$$\frac{\partial f_1}{\partial t} + \vec{v} \cdot \nabla f_1 - \frac{e}{m_e} \vec{E} \cdot \nabla_{\vec{v}} f_0 = 0$$

and

$$\nabla^2 \phi = \frac{e}{\epsilon_0} \int f_1 d^3 \vec{v}$$

Vlasov's solution

- Assume all perturbed quantities vary as $\exp\{i(\vec{k}\cdot\vec{r} - \omega t)\}$
- Solve for f_1 from

$$\frac{\partial f_1}{\partial t} + \vec{v}\cdot\nabla f_1 - \frac{e}{m_e}\vec{E}\cdot\nabla_v f_0 = 0$$

and substitute in

$$\nabla^2\phi = \frac{e}{\epsilon_0} \int f_1 d^3\vec{v}$$

- We get (ϕ has to be non-zero)

$$\frac{e^2}{\epsilon_0 m_e k^2} \int \frac{\vec{k}\cdot\nabla_v f_0}{\omega - \vec{k}\cdot\vec{v}} d^3\vec{v} = -1$$

Dispersion relation

- We got the dispersion relation

$$\frac{e^2}{\epsilon_0 m_e k^2} \int \frac{\vec{k} \cdot \nabla_{\vec{v}} f_0}{\omega - \vec{k} \cdot \vec{v}} d^3 \vec{v} = -1$$

- But there is a singularity at $\omega = \vec{k} \cdot \vec{v}$
- Landau pointed out this problem
- His solution: do not take time dependence as $\exp\{-i\omega t\}$, treat it as an initial value problem:
- f_1 is given at $t = 0$ and we have to find it out at later times.

Landau's solution

Take:

$$f_1(\vec{r}, \vec{v}, t) = f_1(\vec{v}, t) \exp\{i\vec{k} \cdot \vec{r}\}$$

$$\vec{u} = \frac{\vec{k} \cdot \vec{v}}{k} \hat{k}$$

$$\vec{u}_p = \vec{v} - \vec{u}$$

$$F_0(u) = \int f_0(\vec{v}) d^3 \vec{u}_p$$

$$F_1(u, t) = \int f_1(\vec{v}, t) d^3 \vec{u}_p$$

Landau's solution

- Start from

$$\frac{\partial f_1}{\partial t} + \vec{v} \cdot \nabla f_1 - \frac{e}{m_e} \vec{E} \cdot \nabla_v f_0 = 0$$

and

$$\nabla^2 \phi = \frac{e}{\epsilon_0} \int f_1 d^3 \vec{v}$$

- Use Laplace transforms of F_1 and E :

$$\bar{F}_1(u, p) = \int_0^\infty F_1(u, t) e^{-pt} dt$$

$$\bar{E}(p) = \int_0^\infty E(t) e^{-pt} dt$$

Landau's solution

- Some algebra later...

$$\bar{E} = -\frac{e/\epsilon_0}{ik\epsilon(k, p)} \int_{-\infty}^{\infty} \frac{F_1(u, t=0)}{p + iku} du$$

$$\epsilon(k, p) = 1 + \frac{e^2}{\epsilon_0 m_e k} \int_{-\infty}^{\infty} \frac{\partial F_0 / \partial u}{ip - ku} du$$

and

$$\bar{F}_1 = \frac{e}{m_e} \bar{E} \frac{\partial F_0 / \partial u}{p + iku} + \frac{F_1(u, t=0)}{p + iku}$$

Landau's solution

- $\epsilon(k, p)$ is known as the *plasma dielectric function*
- Note that the integrals are well defined
- Now we have to invert these Laplace transforms:

$$a(u, t) = \frac{1}{2\pi i} \int_C \bar{a}(u, p) e^{pt} dp$$

where C is the *Bromwich contour*, which is the contour running parallel to the imaginary axis, and lying to the right of all singularities of \bar{F}_1 in the complex- p plane.

Landau's solution

- Deformation of contours possible when $F_0(u)$ and $F_1(u, t)$ vary smoothly with u .

$$E = \frac{1}{2\pi i} \int_C \left[\frac{-e/\epsilon_0}{ik\epsilon(k, p)} \int_{-\infty}^{\infty} \frac{F_1(u, t=0)}{p + iku} du \right] e^{\rho t} dp$$

E is dominated by the zeros of $\epsilon(k, p)$.

$$F_1 = \frac{1}{2\pi i} \int_C \left[\frac{e}{m_e} \bar{E} \frac{\partial F_0 / \partial u}{p + iku} + \frac{F_1(u, t=0)}{p + iku} \right] e^{\rho t} dp$$

Same with F_1

Landau's solution

- Behaviour depends almost exclusively on the zeros of $\epsilon(k, p)$

$$\epsilon(k, p) = 1 + \frac{e^2}{\epsilon_0 m_e k} \int_{-\infty}^{\infty} \frac{\partial F_0 / \partial u}{ip - ku} du = 0$$

- If $p = -i\omega$, this equation is the same as what we got initially

$$\frac{e^2}{\epsilon_0 m_e k^2} \int \frac{\vec{k} \cdot \nabla_v f_0}{\omega - \vec{k} \cdot \vec{v}} d^3 \vec{v} = -1$$

Dispersion relation

- Use Maxwellian distribution for F_0

$$F_0(u) = \frac{n}{\sqrt{2\pi T_e/m_e}} \exp\left\{-\frac{m_e u^2}{2T_e}\right\}$$

- Taylor-expanding denominator

$$\frac{1}{\omega - ku} = \frac{1}{\omega} \left(1 + \frac{ku}{\omega} + \frac{k^2 u^2}{\omega^2} + \dots\right)$$

- Treat imaginary part of frequency as a perturbation

$$\omega \approx \Re\{\omega\} = \omega_0, \quad \omega = \omega_0 + \delta\omega$$

Dispersion relation

After some complex analysis and algebra, we finally get

$$\omega^2 \approx \omega_p^2(1 + 3k^2\lambda_D^2)$$

and

$$\begin{aligned}\delta\omega &\approx \frac{i\pi}{2} \frac{e^2\omega_p}{\epsilon_0 m_e k^2} \left(\frac{\partial F_0}{\partial u}\right)_{u=\omega/k} \\ &\approx \frac{-i}{2} \sqrt{\frac{\pi}{2}} \frac{\omega_p}{(k\lambda_D)^3} \exp\left\{\frac{-1}{2(k\lambda_D)^2} - \frac{3}{2}\right\}\end{aligned}$$

where λ_D is the Debye length and ω_p is the plasma frequency

$$\lambda_D = \sqrt{\frac{T_e}{m_e \omega_p^2}} \quad \text{and} \quad \omega_p = \sqrt{\frac{ne^2}{\epsilon_0 m_e}}$$

Landau damping

- Damping happens due to the imaginary term

$$\delta\omega \approx \frac{-i}{2} \sqrt{\frac{\pi}{2}} \frac{\omega_p}{(k\lambda_D)^3} \exp\left\{\frac{-1}{2(k\lambda_D)^2} - \frac{3}{2}\right\}$$

This is the *Landau damping*.

- Large for shorter wavelengths, negligible for longer wavelengths.

Landau damping: Verification

- Equation of motion of a charged particle:

$$\frac{d^2x}{dt^2} = \frac{e}{m} E_0 \exp\{i(kx - \omega t)\}$$

- Taking $x = x_0 + u_0 t$ in the electric field, we get

$$u - u_0 = \frac{e^{ikx_0} e E_0}{i(ku_0 - \omega)m} [\exp\{i(ku_0 - \omega)t\} - 1]$$

Landau damping: Verification

- Particles far from $u_0 = \omega/k$ will simply oscillate.
- If $u_0 \approx \omega/k$, then

$$u - u_0 = \frac{eE_0}{m} t e^{ikx_0}$$

These are *resonant* particles, some of them gain energy from the wave, and some lose energy to the wave.

Landau damping: Verification

- For a particle with u_0 slightly greater than ω/k
 - loses energy \implies comes close to ω/k (more interaction)
 - gains energy \implies goes away from ω/k
- For a particle with u_0 slightly lesser than ω/k
 - loses energy \implies it goes away from ω/k
 - gains energy \implies it comes close to ω/k (more interaction)
- On average
 - particles with $u_0 \gtrsim \omega/k$ lose energy
 - particles with $u_0 \lesssim \omega/k$ gain energy

Landau damping

- Intuition: Surfers on a wave
- If more particles have $u_0 \lesssim \omega/k$ than particles with $u_0 \gtrsim \omega/k$:
 - \implies net transfer of energy from \vec{E} to the particles
 - i.e., the electric field is damped.
- Similar effect in galactic dynamics:
 - Electrons \leftrightarrow stars, electric field \leftrightarrow gravity
 - Vlasov eqn. is called *Collisionless Boltzmann eqn.*

References

- Plasma Physics web notes, R. Fitzpatrick, U. Texas Austin

<http://farside.ph.utexas.edu/teaching/plasma/Plasmahtml/>

- Wikipedia

https://en.wikipedia.org/wiki/Landau_damping